

Herzberg Institute of Astrophysics

CASTOR

Cosmological Advanced Telescope for Optical & UV Research

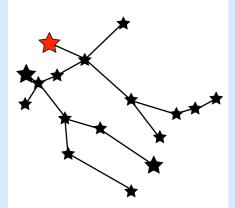
National Research Council Canada Conseil national de recherches Canada



NRC · CNRC

Herzberg Institute of Astrophysics

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Cosmological Advanced Telescope for Optical & UV Research



Conseil national de recherches Canada

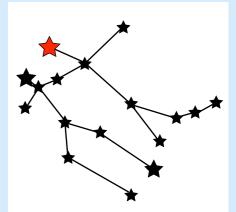


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Conseil national de recherches Canada



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Canadian Led

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Canadian LedWide Field

anao

Canadian Led
Wide Field
UV-Optical Imaging

anao

Canadian Led
Wide Field
UV-Optical Imaging
Survey Optimized

anao

Canadian Led
Wide Field
UV-Optical Imaging
Survey Optimized
Space Telescope

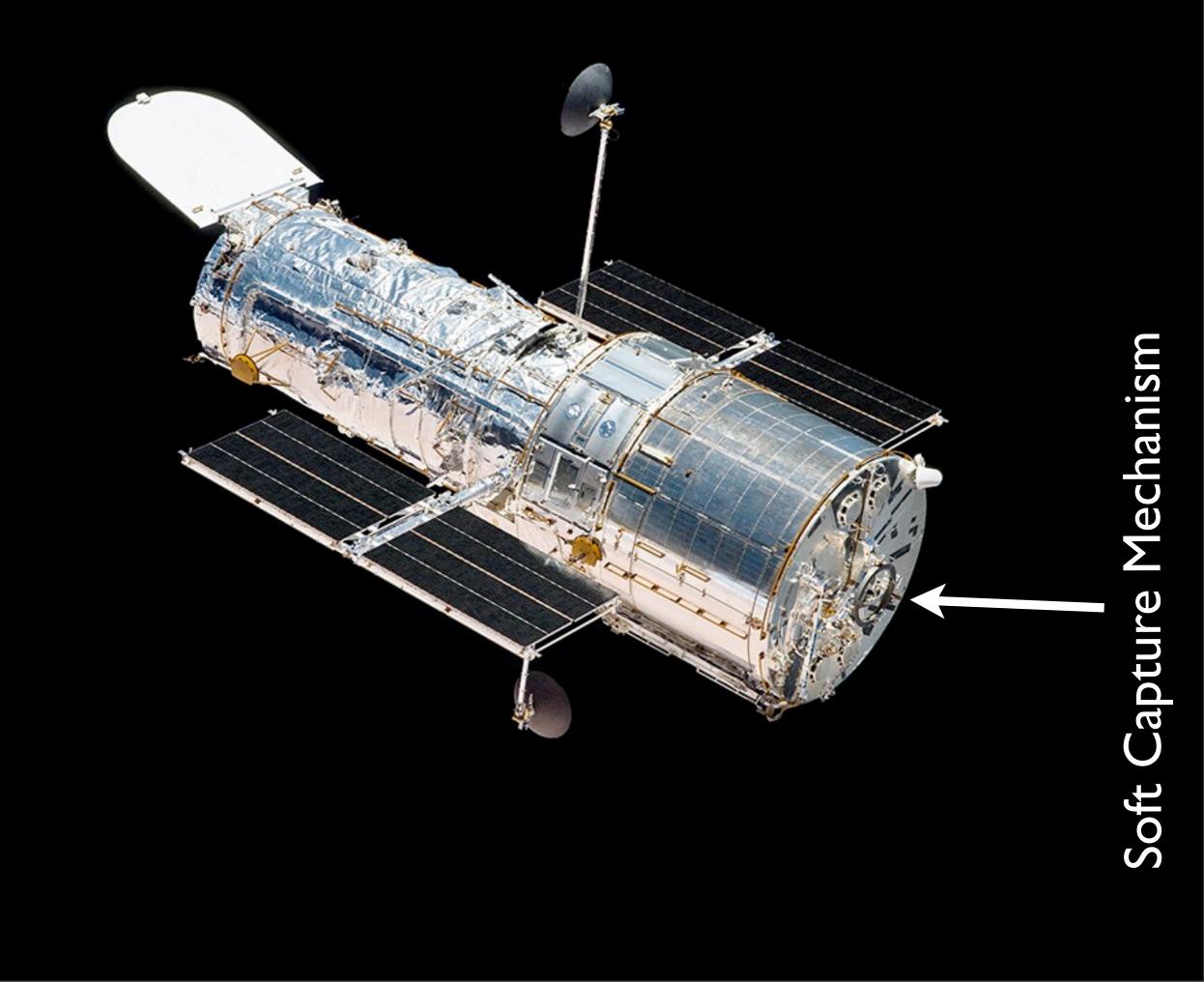


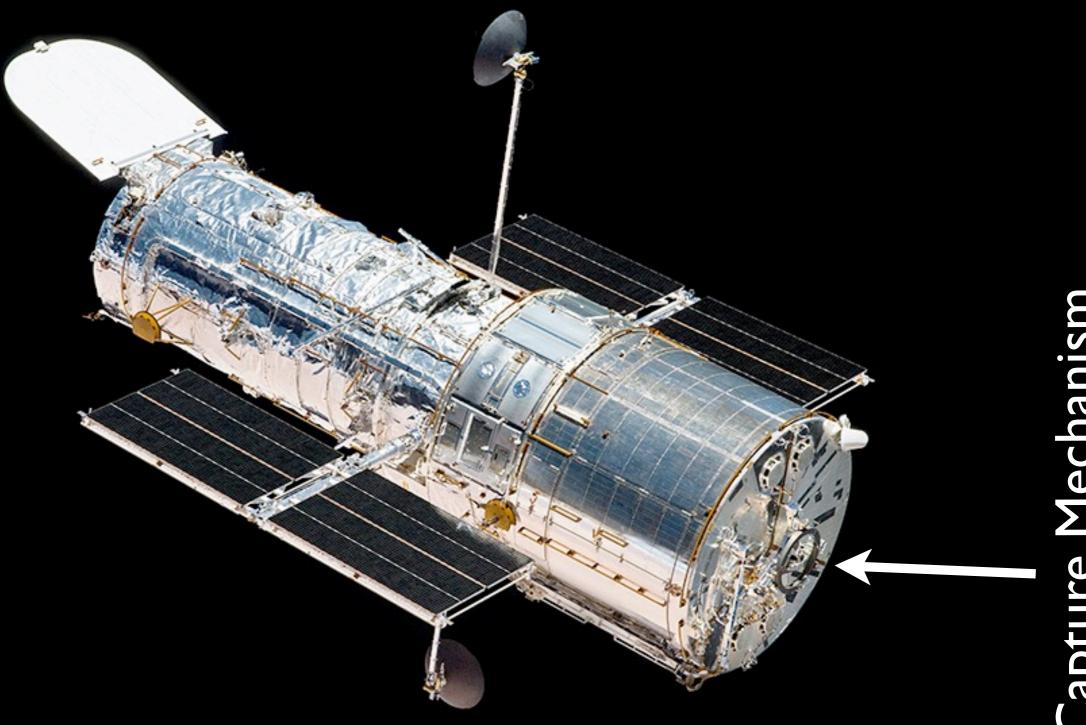
Sombrero Galaxy • M104



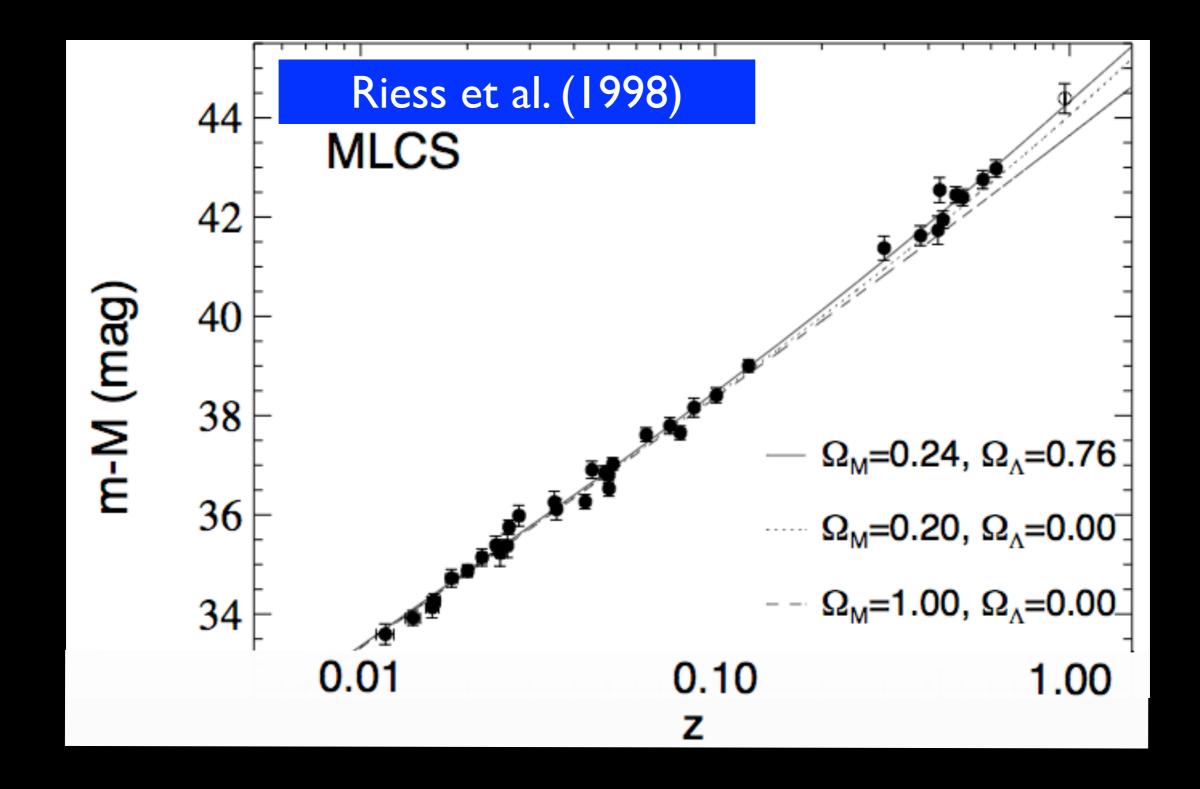


NASA and The Hubble Heritage Team (AURA/STScl) • Hubble Space Telescope ACS • STScl-PRC03-28

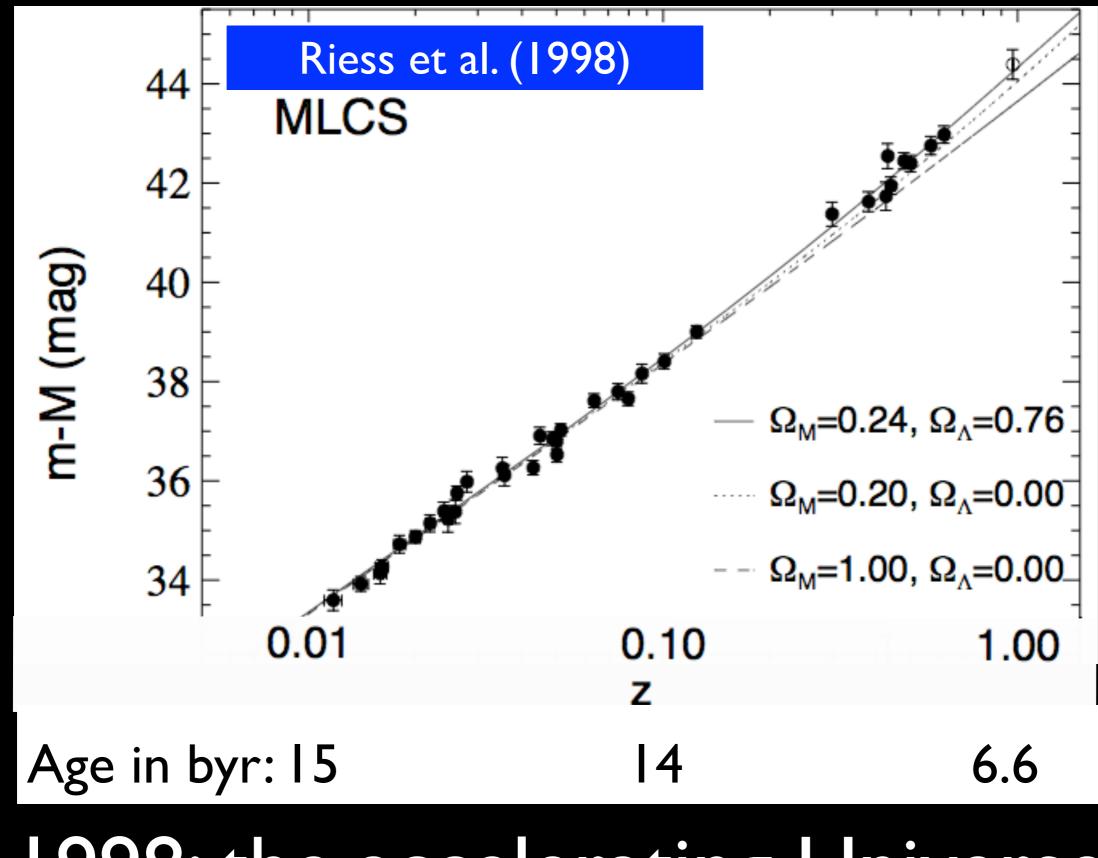




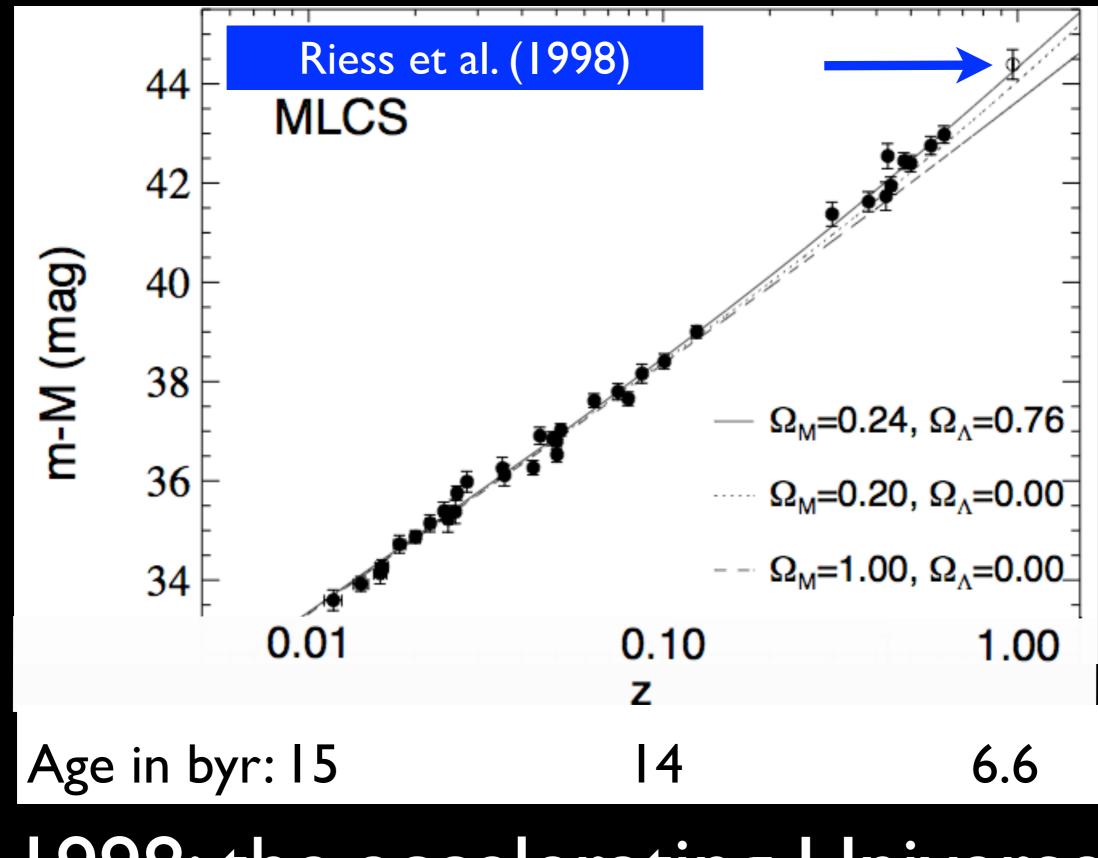
Hubble's Farewell



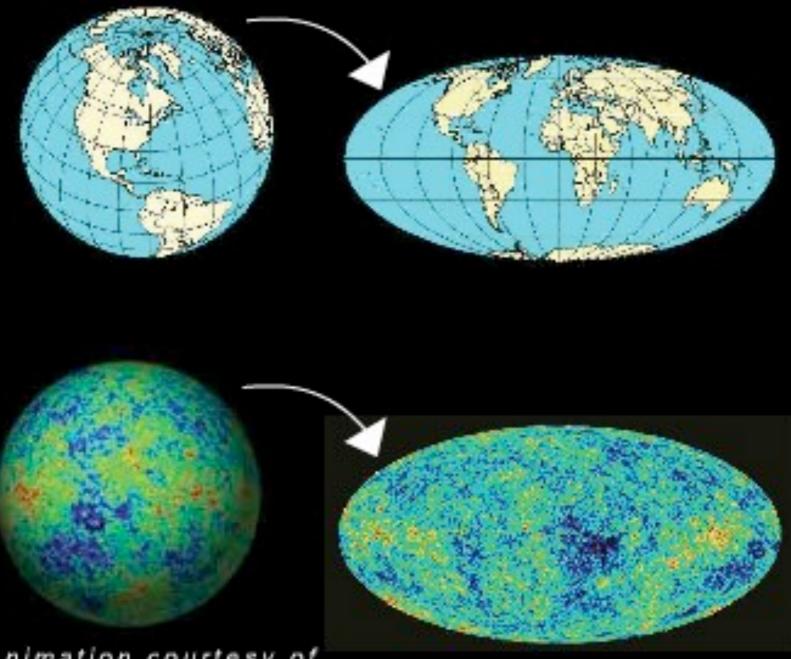
1998: the accelerating Universe



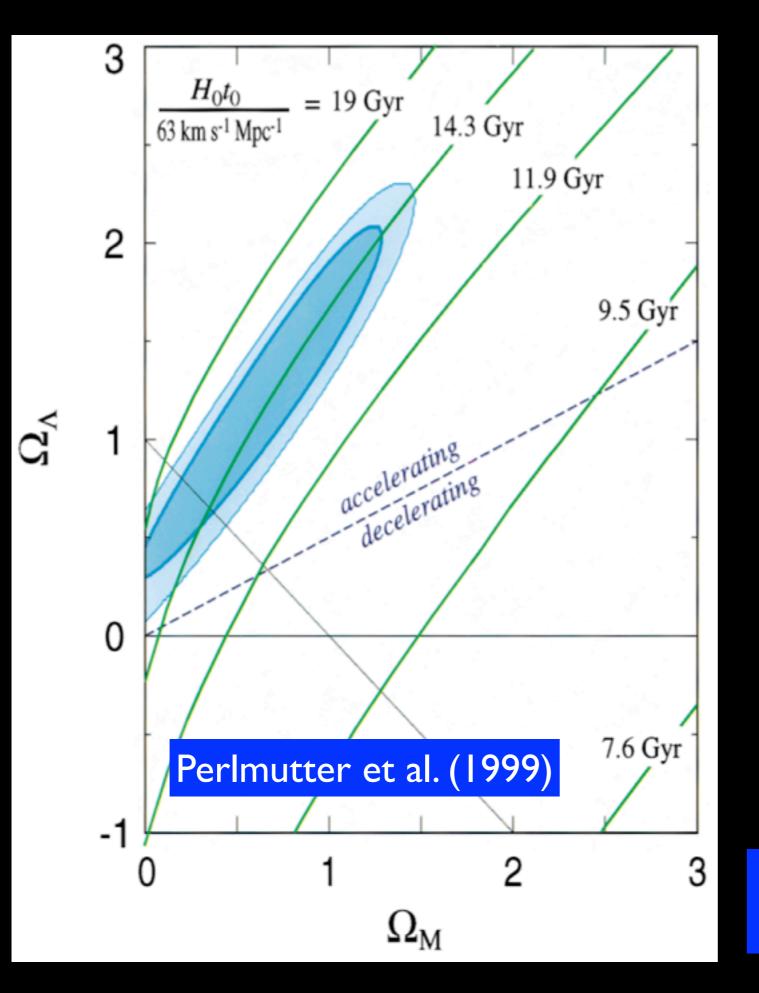
1998: the accelerating Universe



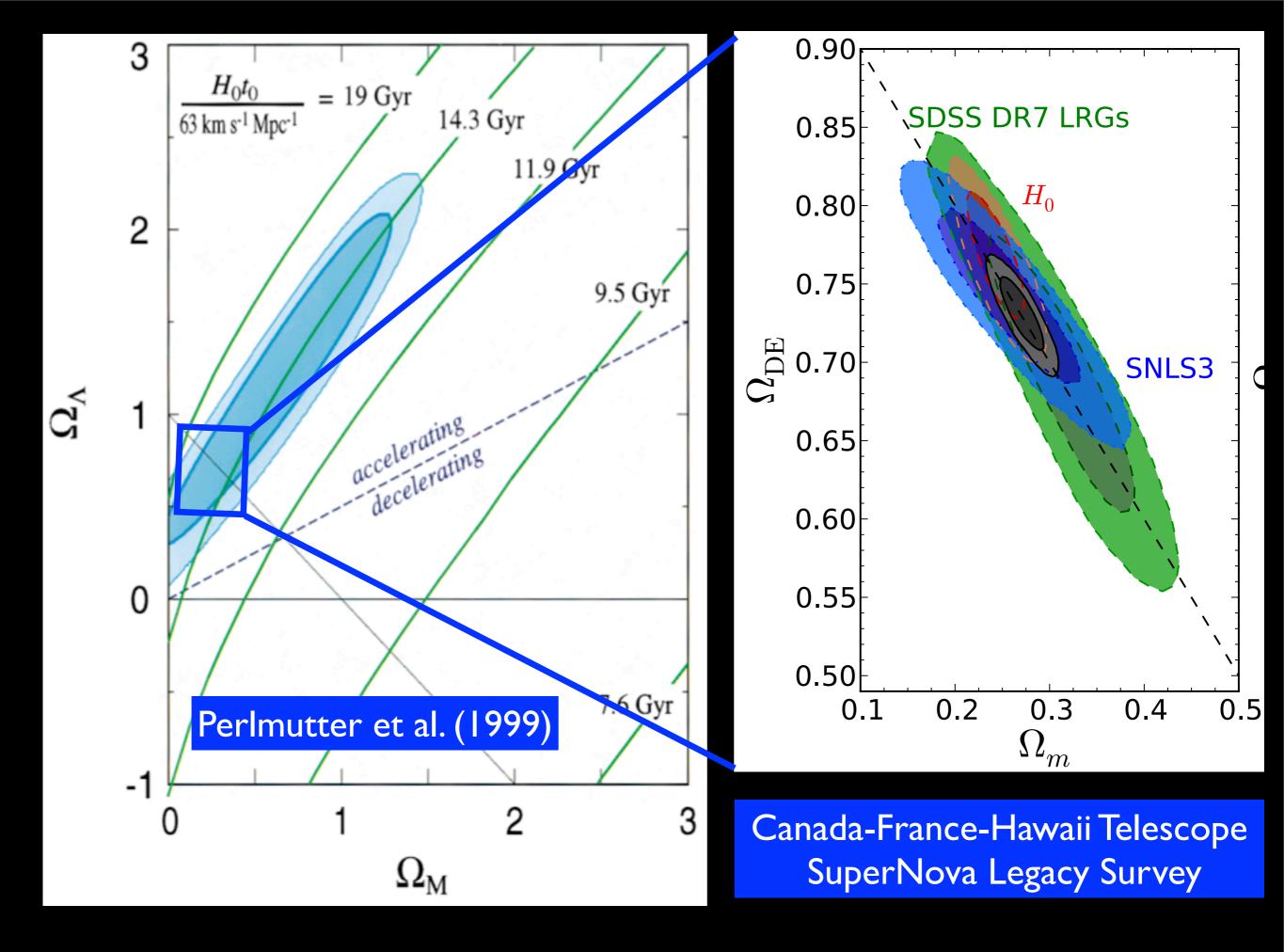
1998: the accelerating Universe

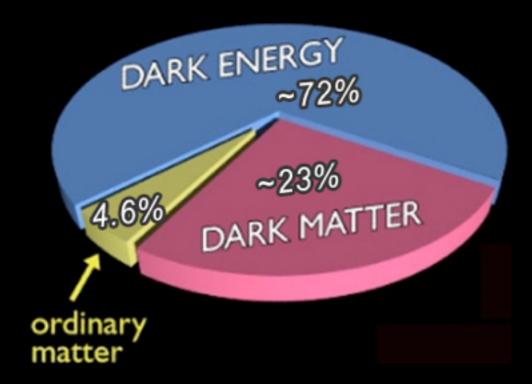


Animation courtesy of NASA and WMAP



Canada-France-Hawaii Telescope SuperNova Legacy Survey



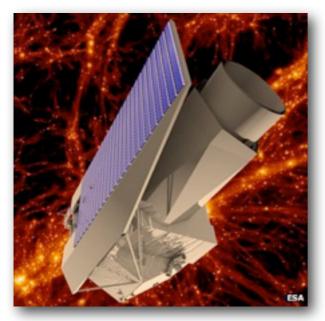




Standard Model

Dark Matter Dark Energy Cosmic Inflation Matter

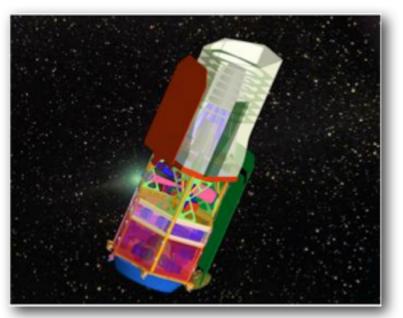
International Plans and Priorities



Euclid (ESA)



LSST (USA)

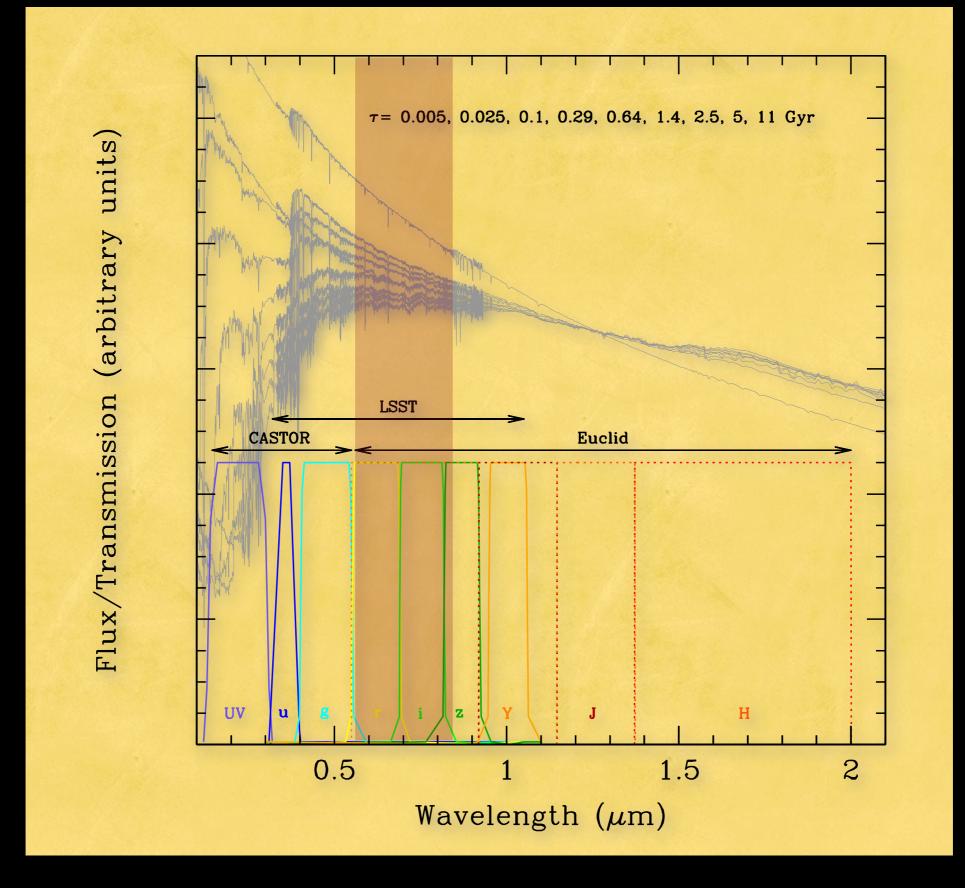


WFIRST (NASA)

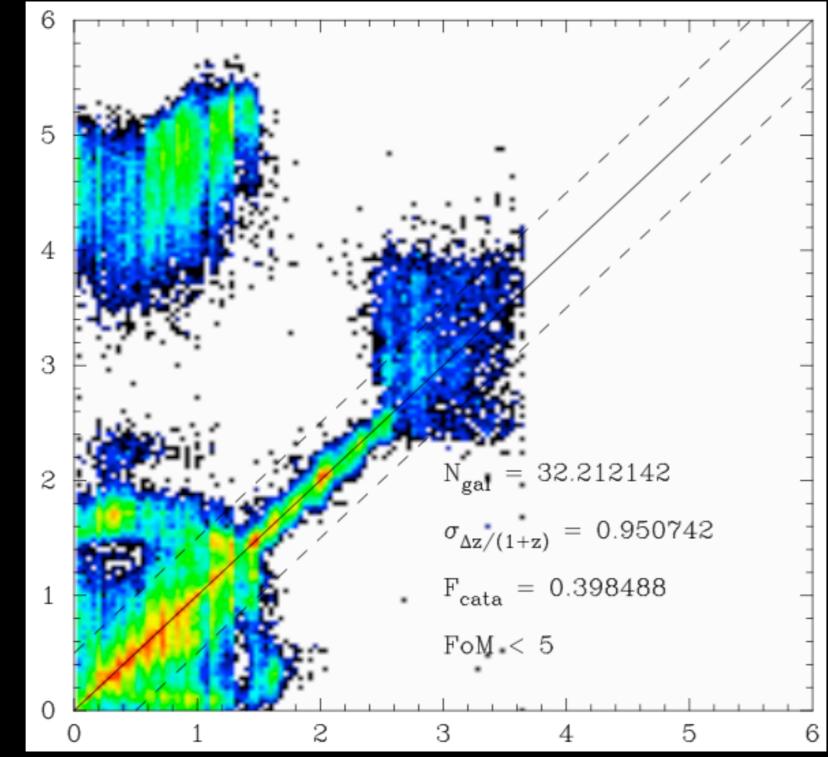
Lead Agency	ESA	
Aperture	1.2m	
Location	Earth-Sun L2 point	
Launch	2018-2019	
Lifetime	6 years	
Depth	24.5 AB mag	
Sky Coverage	15,000 deg ²	
Visible Imager	550 – 900 nm (RIZ)	
IR Imager	930 – 2000 nm (YJH)	
IR Spectroscopy	R ~ 250 (slitless)	

Lead Agency	NSF/DoE	Lead Agency	NASA
Aperture	6.7m (unobscured)	Aperture	1.5m
Location	Cerro Pacon, Chile	Location	Earth-Sun L2 point
Launch	~2019	Launch	~2025:
Lifetime	nominal 10 years	Lifetime	5 years
Depth	26.1 (u), 27.4 (g)	Depth	25.5 AB mag
Sky Coverage	20,000+ deg ²	Sky Coverage	20,000 deg ²
Visible/IR Imager	330 - 1050 nm (ugrizY)	Visible/IR Imager	400 – 2000 nm
IR Spectroscopy	None	IR Spectroscopy	R > 75 (slitless)

• The characterization of dark energy is a primary goal for each of these facilities. No single facility is expected to solve the dark energy mystery — complementarity is essential.



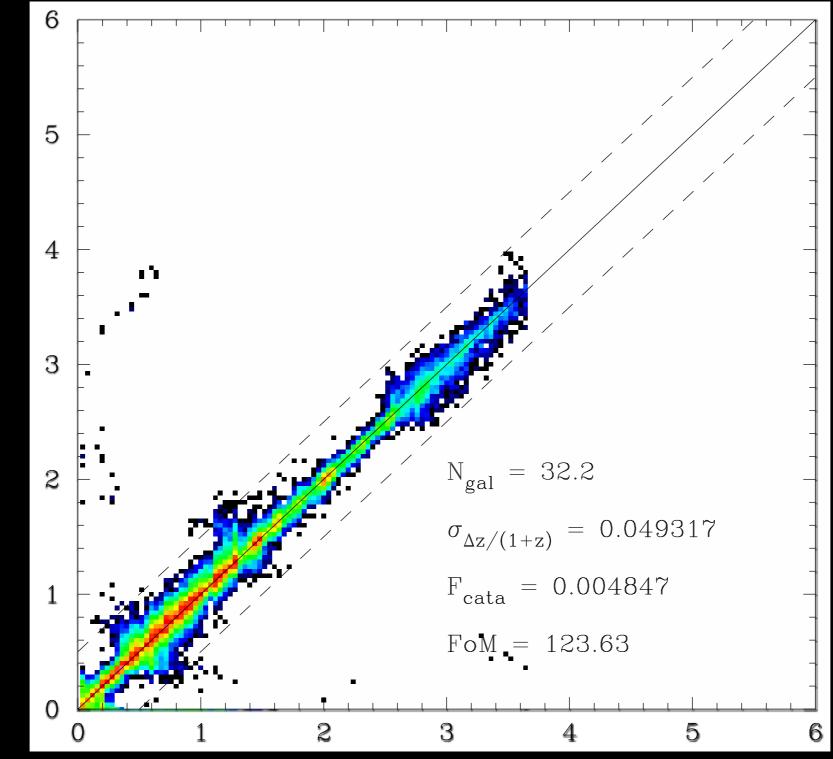
Euclid + CASTOR



Actual

Measure

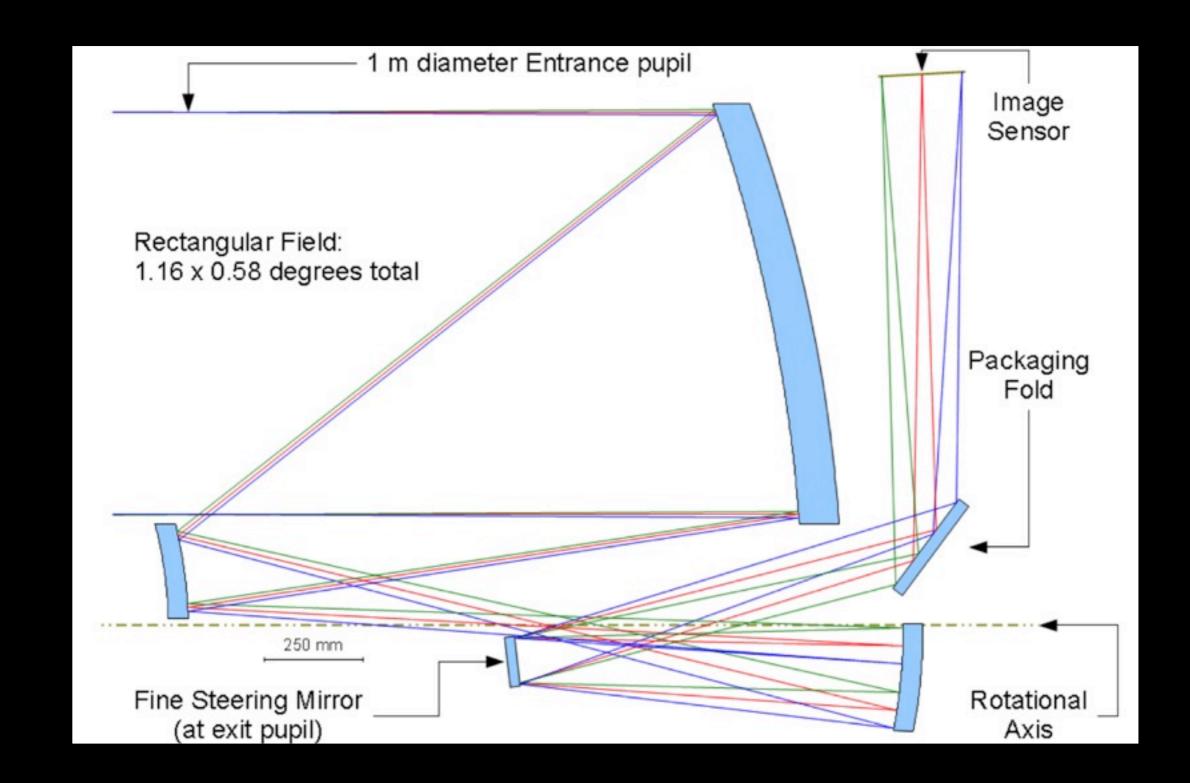
Euclid + CASTOR



Actual

l e

Measu

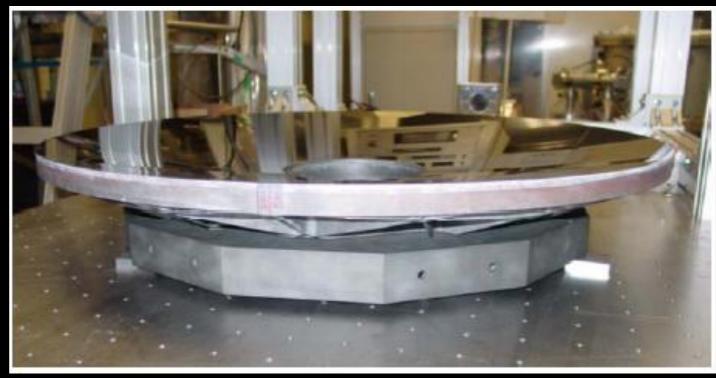




back

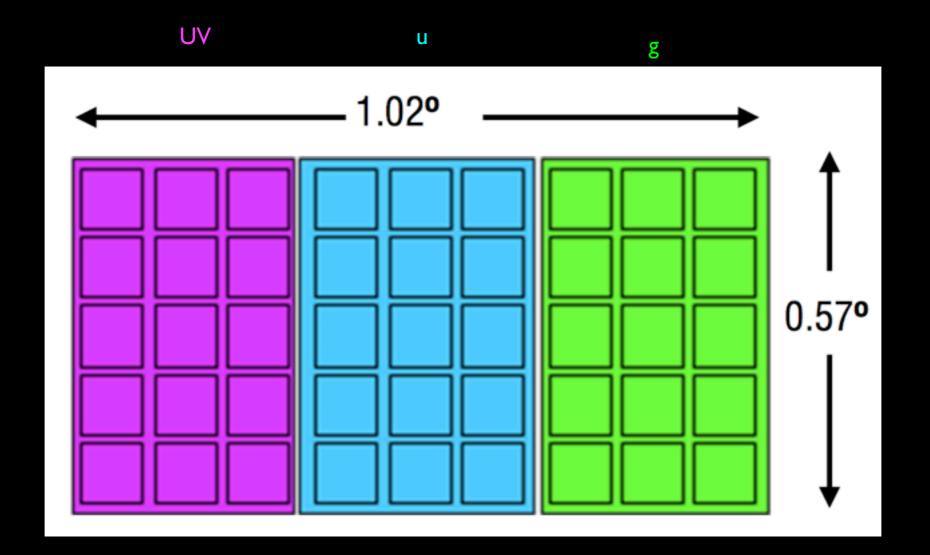
Built for Astro-H

Keep the mass down with light weight SiC mirror



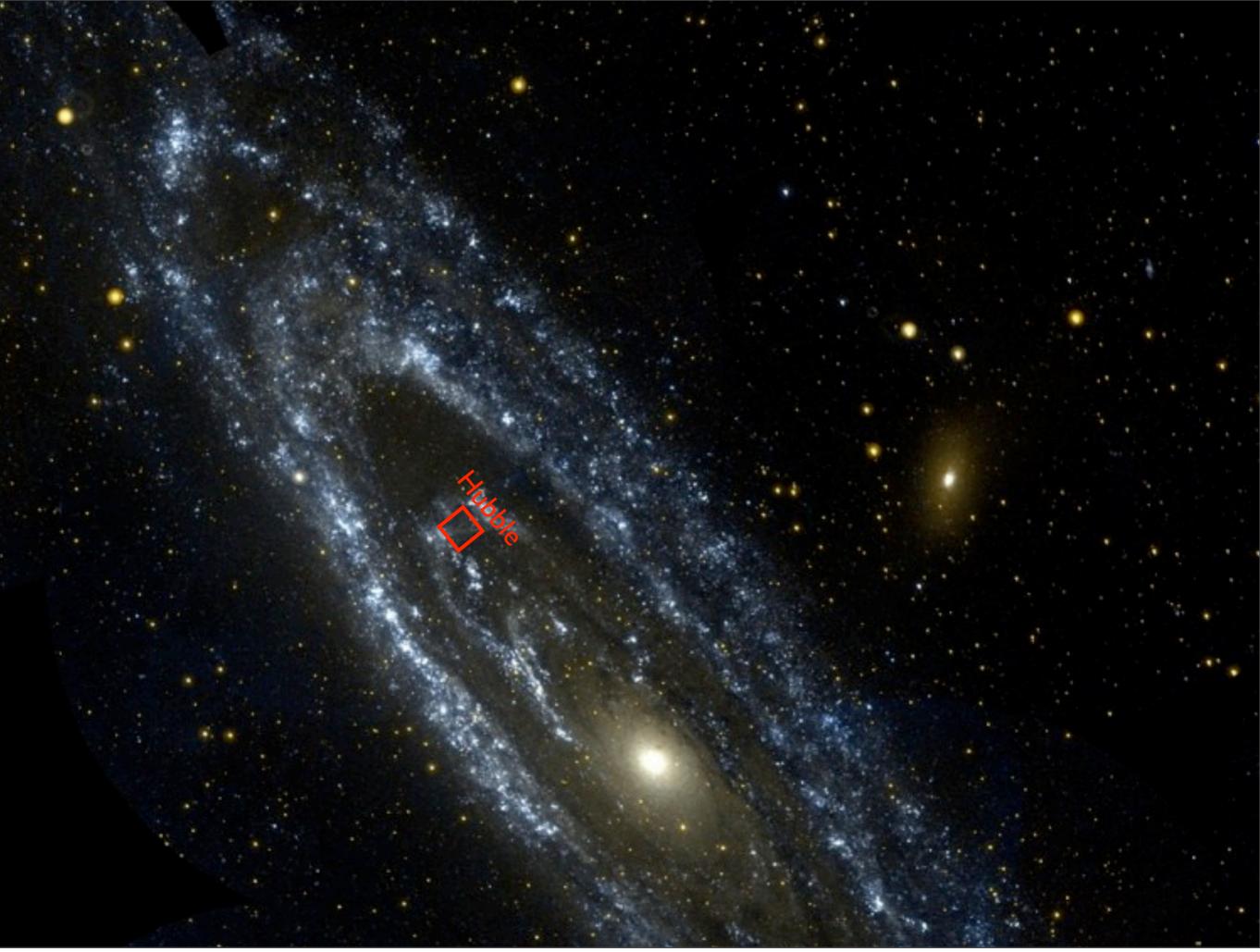
front

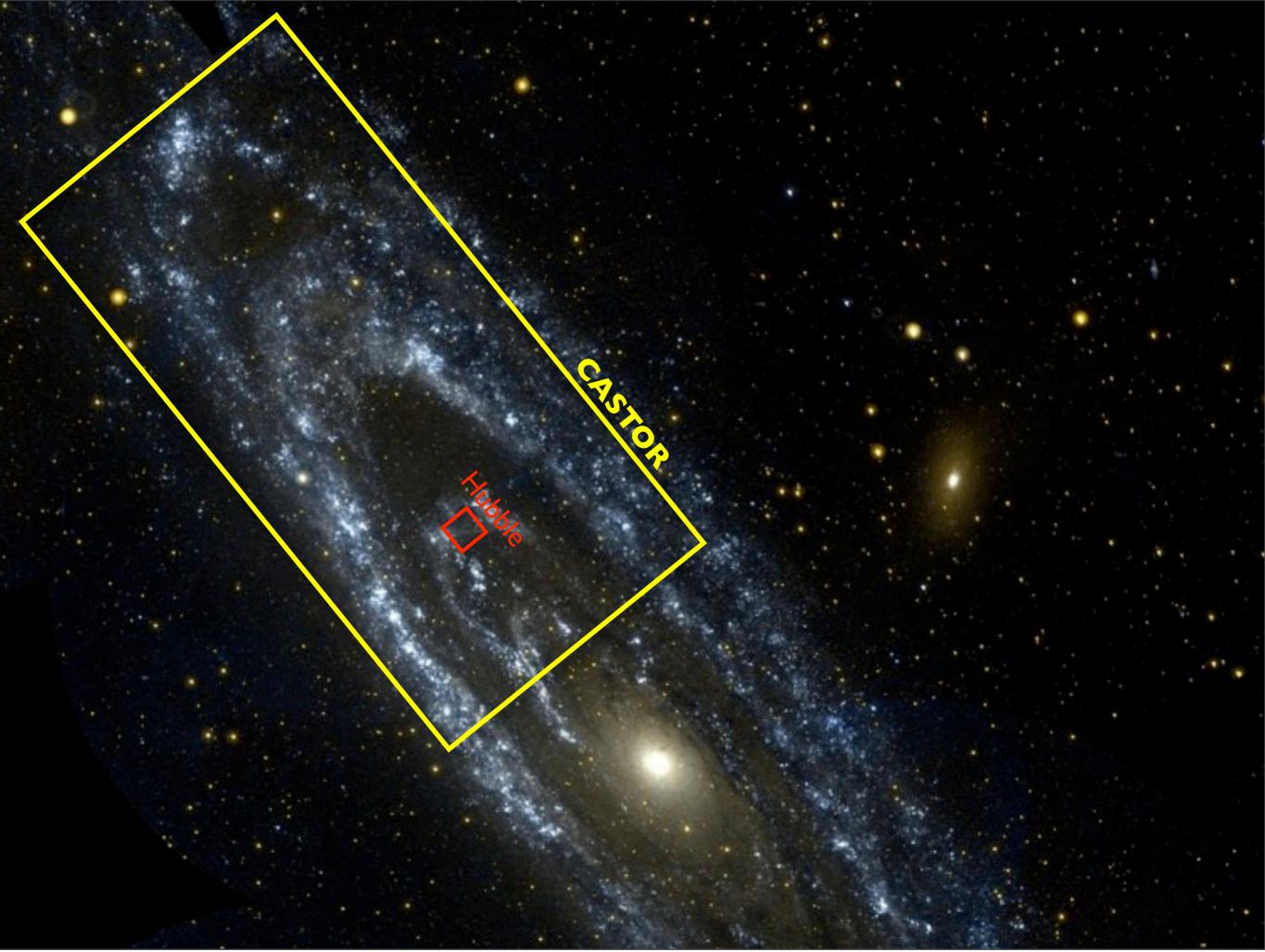
a 'giga pixel' camera in space

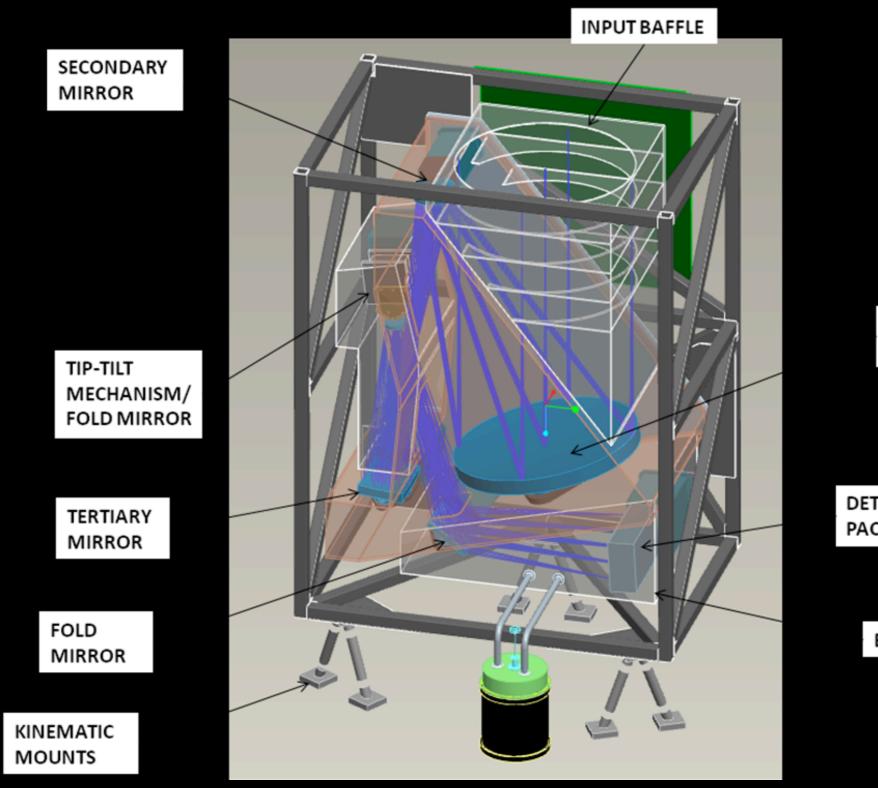


5 x 9 array of 4096x4096 H4RG-10 CMOS detectors Used in James Web and others





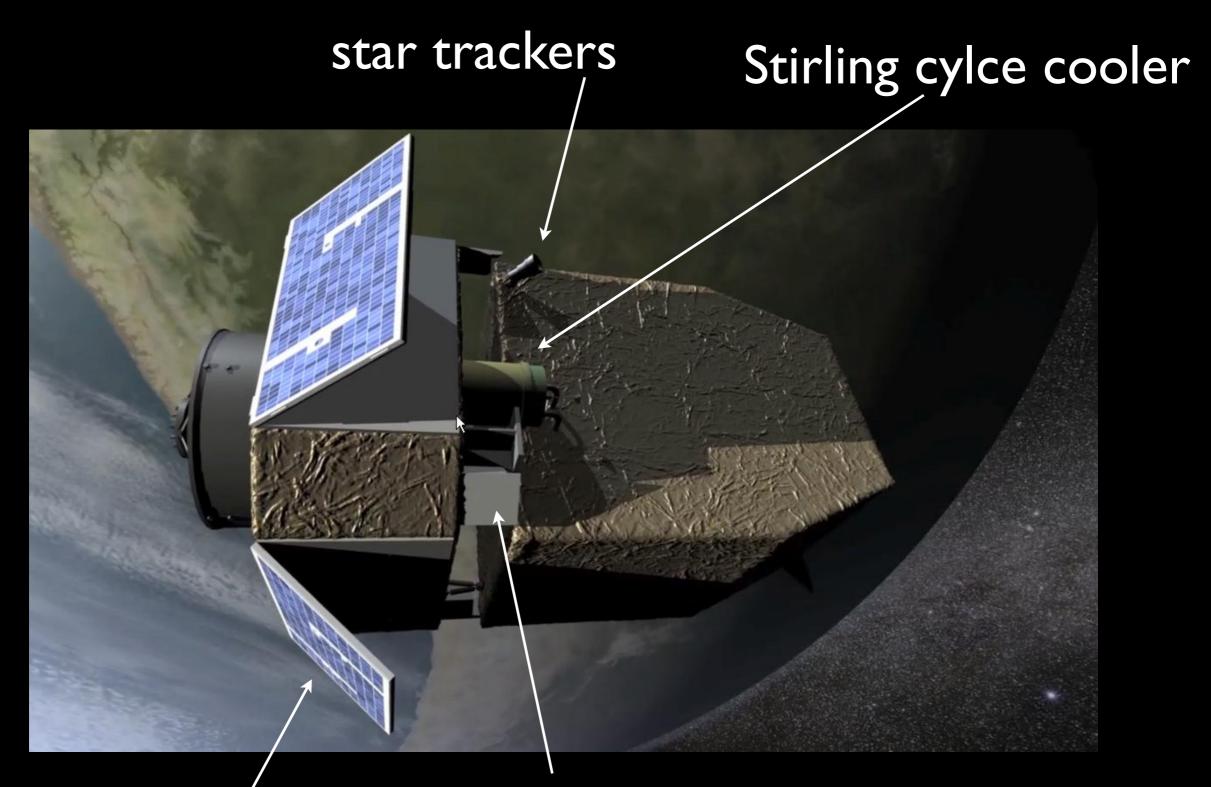




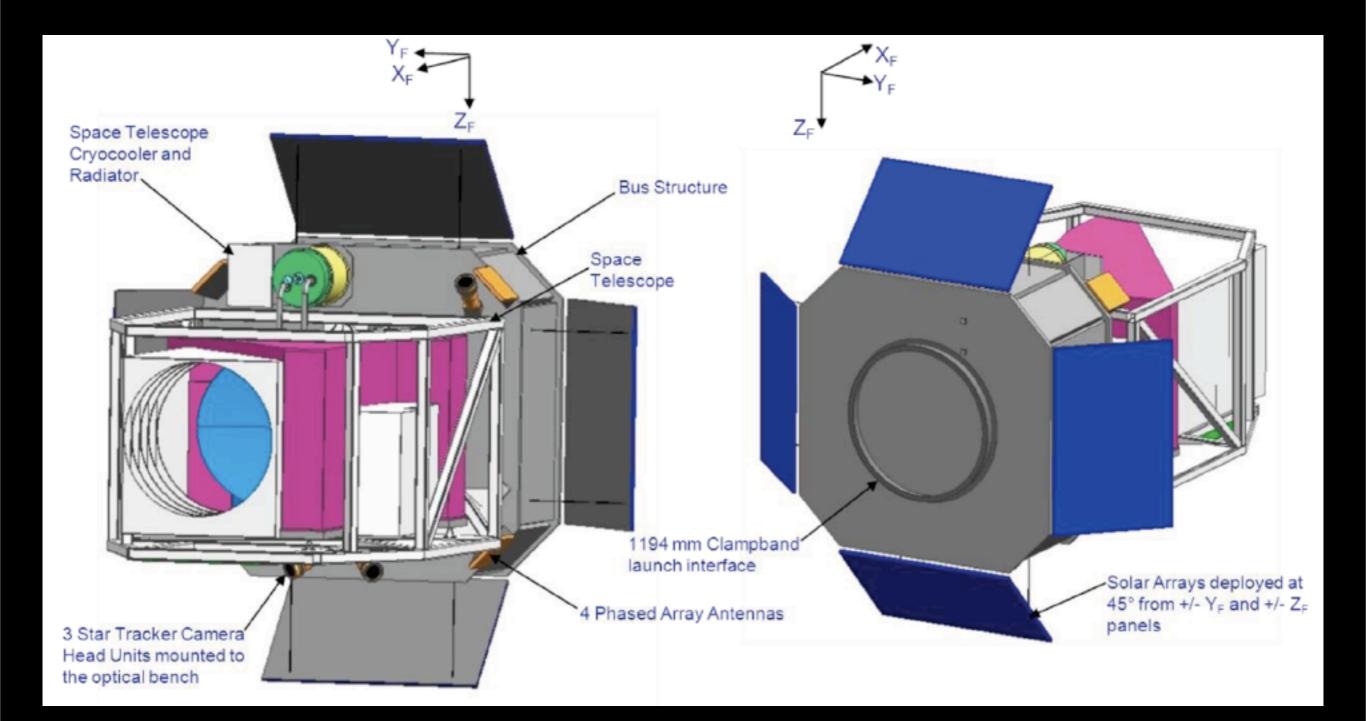
PRIMARY MIRROR

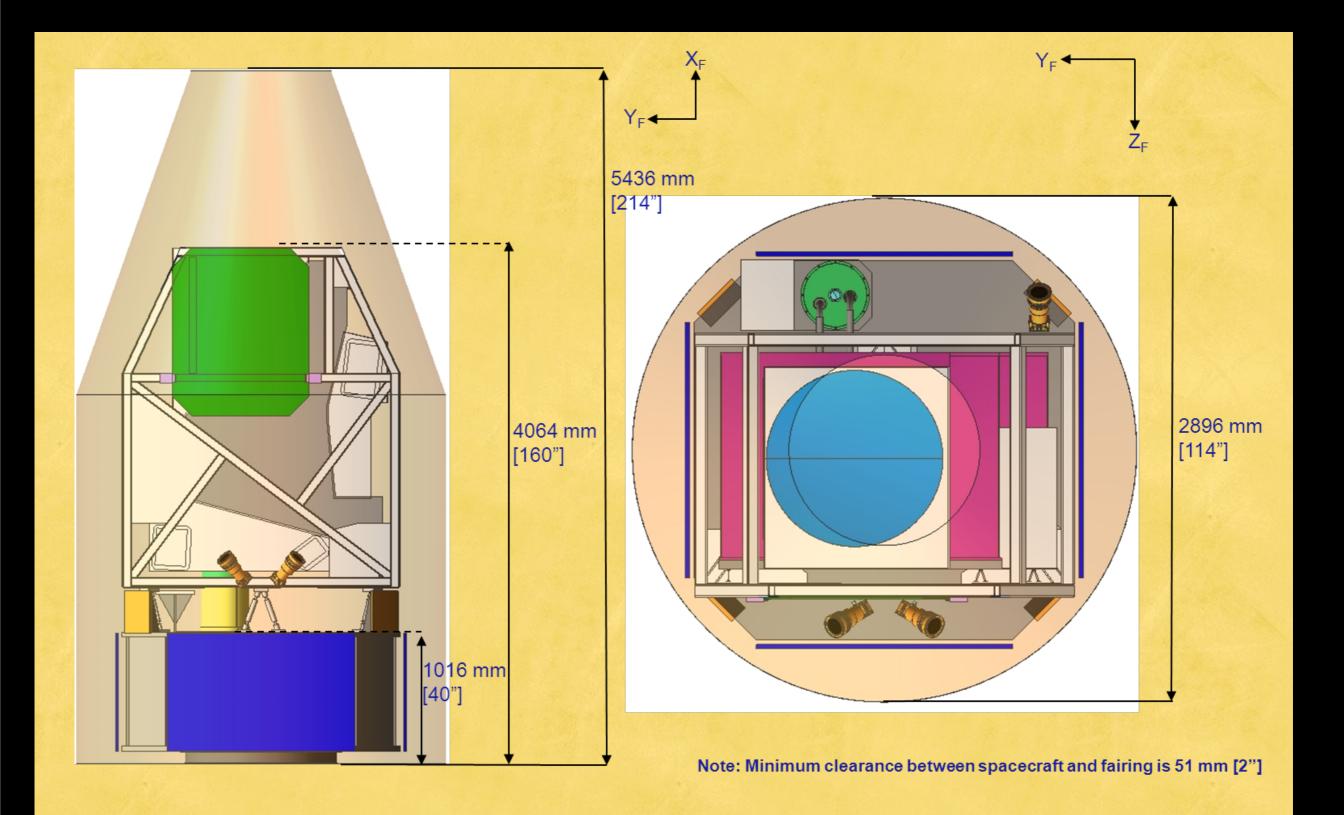
DETECTOR PACKAGE (FPA)

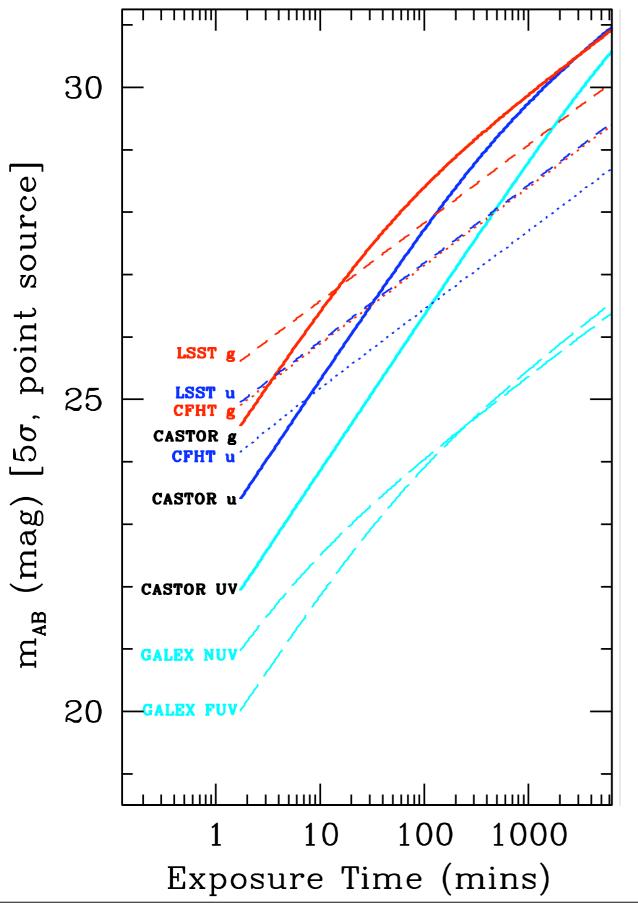
EXIT BAFFLE

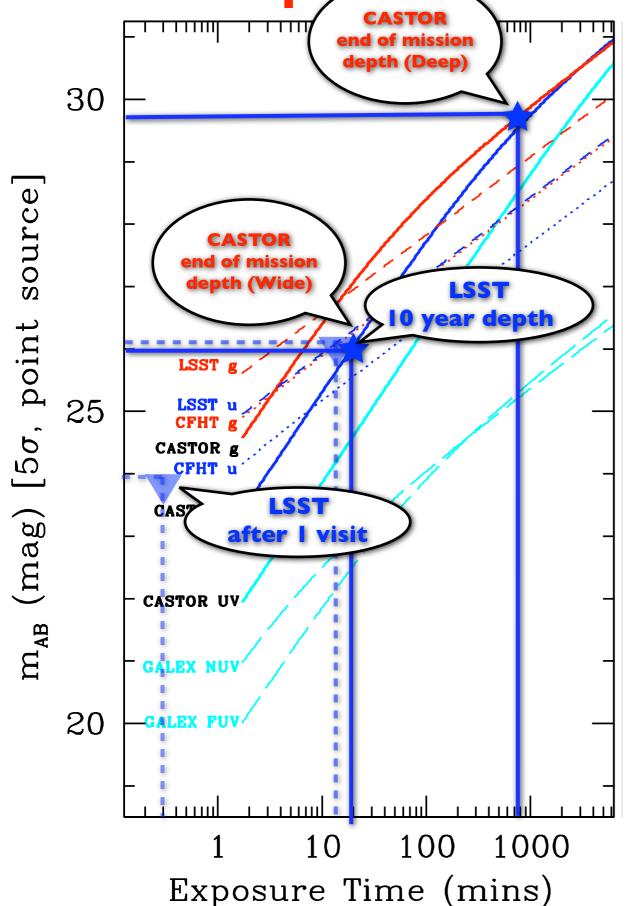


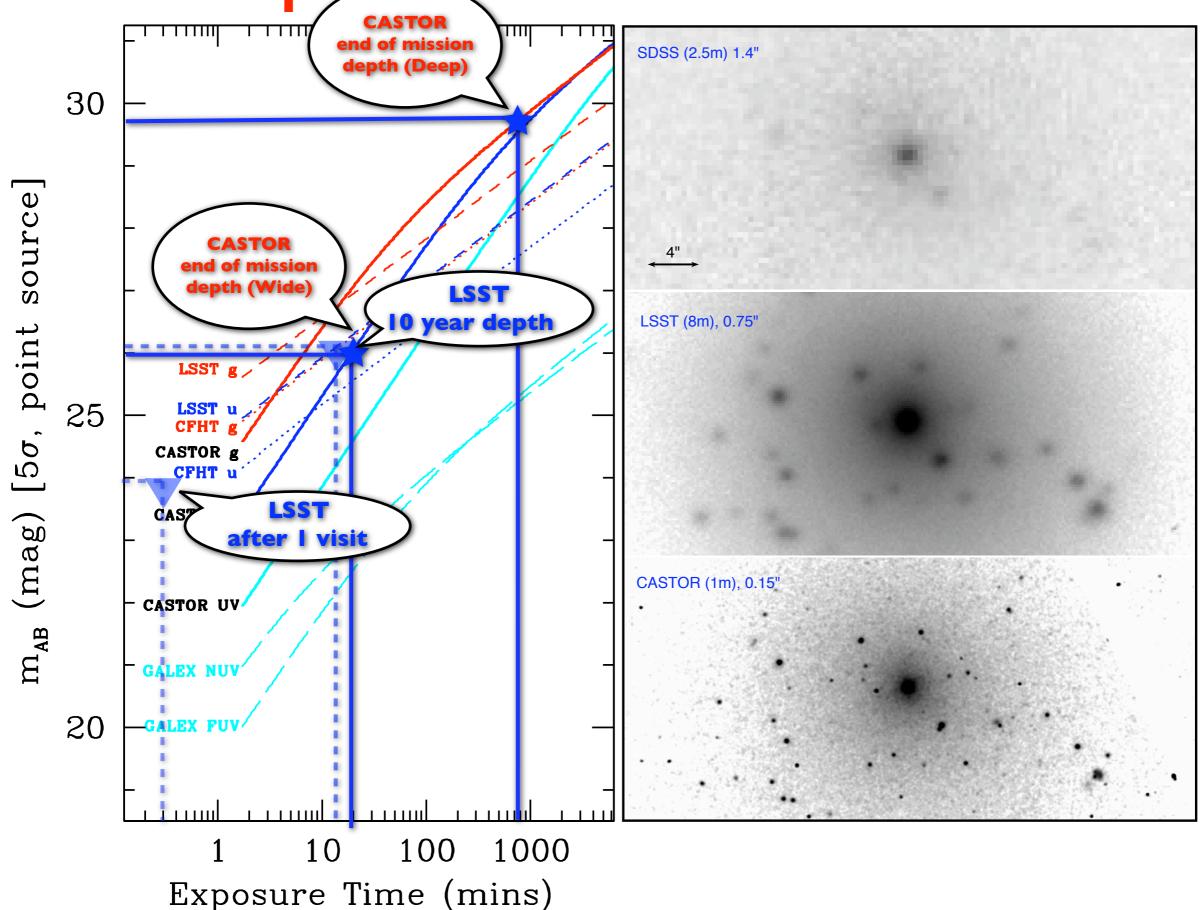
/ X-Band Phased Array Downlinks Deployable solar panels





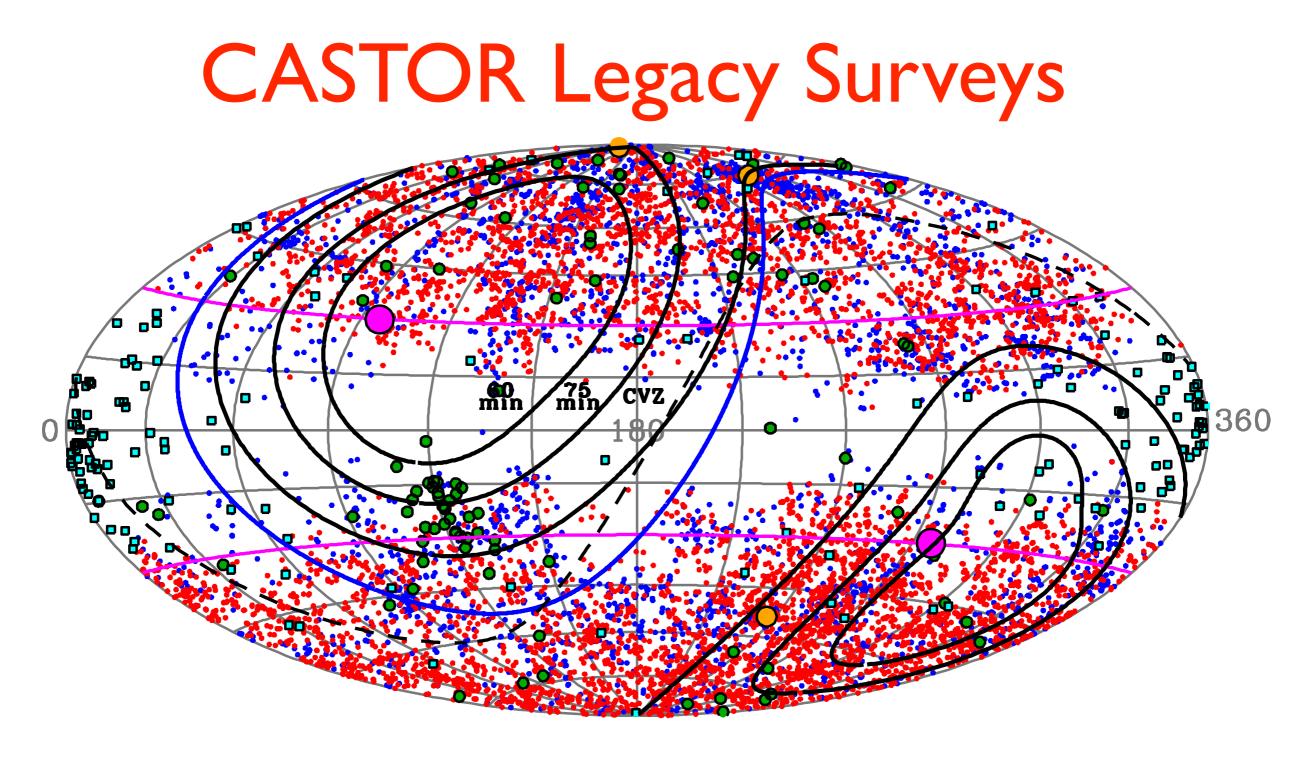






Mission Agency	Duration	Diameter	lmage Quality, θ	Field of View, Ω	Survey Speed Relative to HST
Hubble Space Telescope NASA	1990-2018?	2.5m	0.1″	0.0031 deg ²	Ι
GALEX	2003-2012	0.5m	5″	I.I deg ²	0.14
NASA					
UVIT	2013-2018	2 x 0.38m	I.5″	0.2 deg ²	0.29
ISRO, CSA					
CASTOR	2020+	lm	0.15″	0.67 deg ²	96.1
CSA					

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CSA					96.1



- CVZ, 75min, 60min viewing zones
- — · Ecliptic Plane

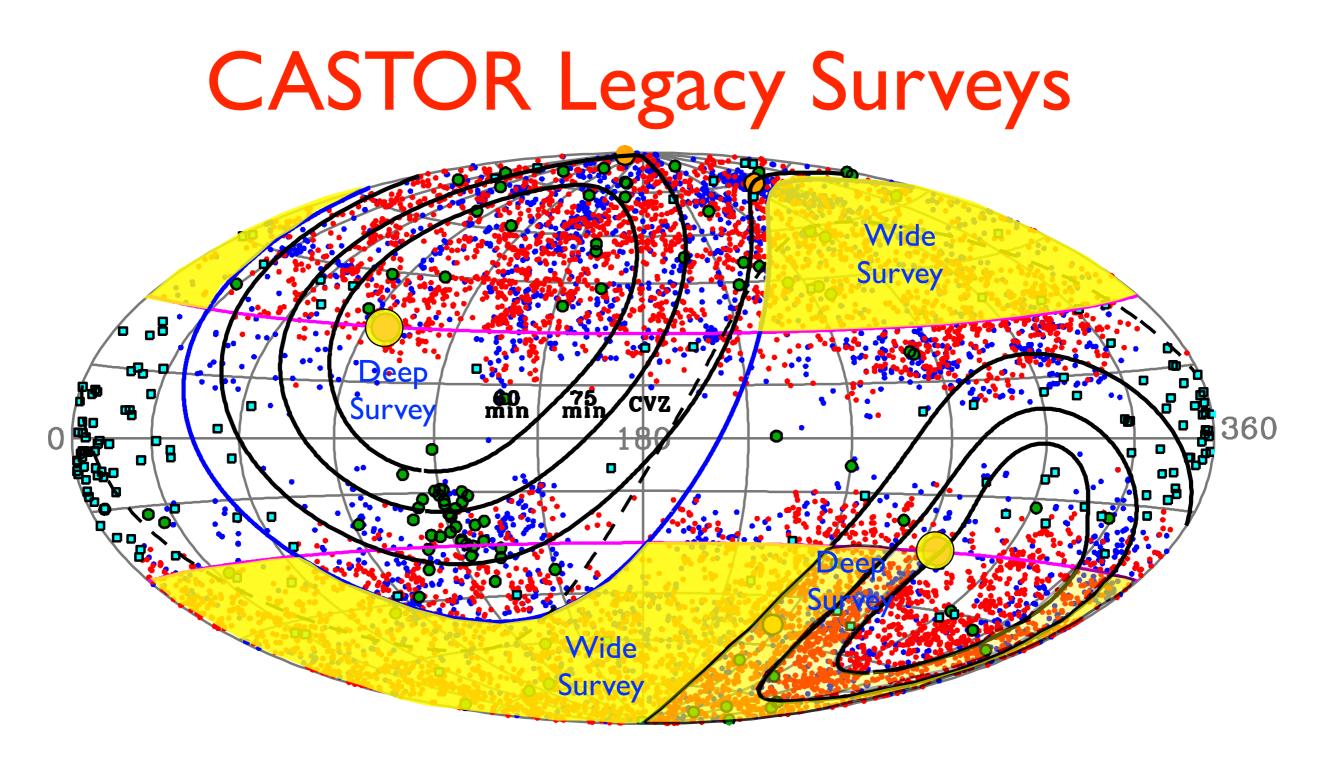
Euclid Wide Survey Limits $(|b| > 30^\circ)$

- **Euclid Deep Fields**
- LSST Survey Limit ($\delta < +10^\circ$)

- Nearby Galaxies (D < 3 Mpc)
- **RC3** Galaxies

0

- Abell Clusters
- Milky Way Globular Clusters
- O Virgo, Fornax, Coma



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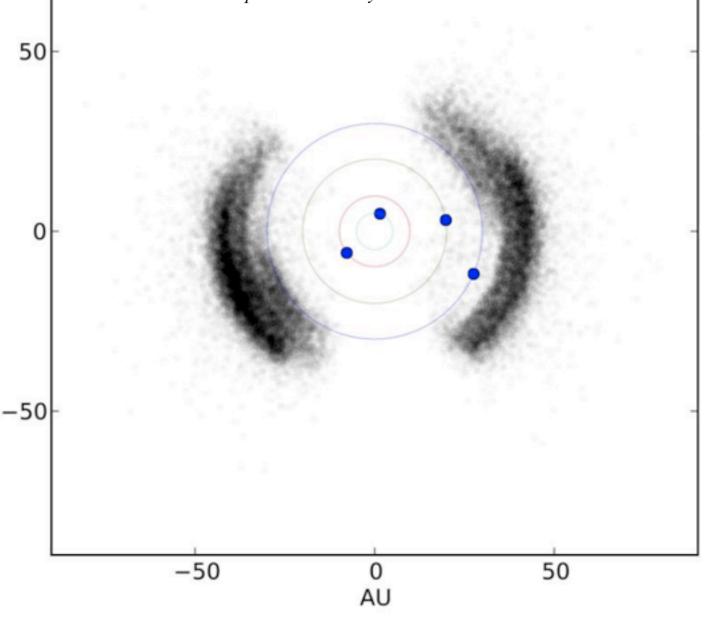
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- Milky Way Globular Clusters
- Virgo, Fornax, Coma

Science: Outer Solar System

AU

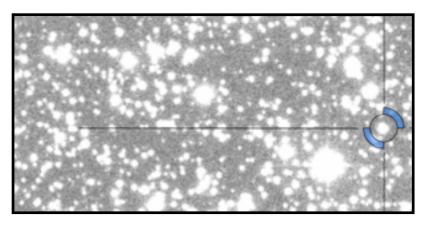
- The outer solar system (OSS) is the region from the orbit of Neptune (30 AU) out to a few thousand AU.
- It contains many components of interest, including the centaur, Kuiper Belt, and scattered disk populations.
- At present, only ~1500 OSS objects have been catalogued.
- Only two OSS objects are known beyond 50 AU.

A top-down view of the outer solar system based on predictions from the Canada-France-Ecliptic Plane Survey model.



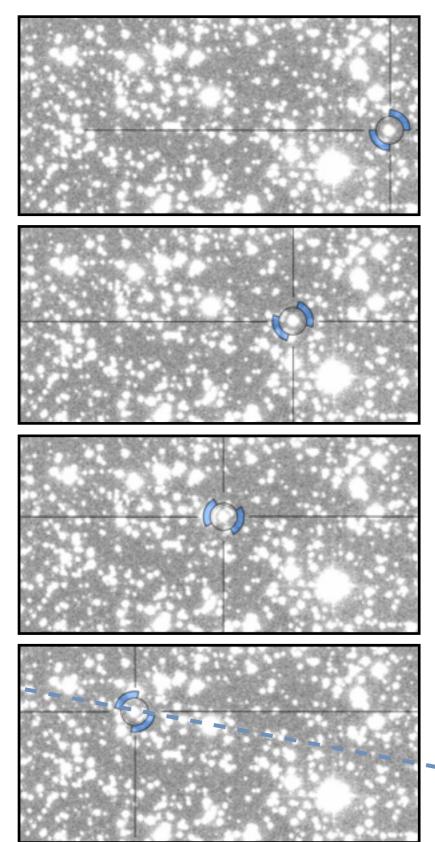
Science: Outer Solar System

- The role of CASTOR:
 - Moving object searches in the CASTOR Wide Survey would discover ~20000 new OSS objects, approximately 50 of which will be beyond 400 AU!
 - Occultation surveys of stars by OSS objects has the potential to detect, and characterize, the population of sub-km sources. This would require a dedicated survey requiring a few weeks or months.
 - Surface chemistry constraints (including the presence of organic ices) for OSS objects are available from their spectral energy distributions, particularly at the shortest wavelengths. The CASTOR Wide Survey would provide u-band photometry for ~800 OSSs (compared to the present sample of 10 with u-band data).



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CASTOR Summary

- CASTOR: Cosmological Advanced Survey Telescope for Optical and UV Research
 - A nearly diffraction-limited 1 m telescope (FWHM = 0.15"), focused on wide-field imaging (> 0.5 deg²) at UV and blue-optical wavelengths (150 550 nm).
- **CASTOR** is a potential *flagship* Canadian space astronomy mission that would:
 - make a significant and strategic contribution to future Dark Energy missions (Euclid, WFIRST, LSST).
 - provide a natural UV/optical successor to the Hubble Space Telescope, with a 200x gain in field of view.
 - fulfill the requirements of the 2010 Long Range Plan for Canadian Astronomy.
 - represent an important next step in the long-term development of the Canadian space program.
 - act as a catalyst for new international collaborations and partnerships.
 - serve as a high-profile showcase for Canadian technological capabilities.

